

Effective field theory in the open systems
Schwinger-Keldysh formalism and its gravity dual

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0. Outline

- Closed and open systems of IR modes
- Schwinger-Keldysh formalism (CTP)
- Generating functional and effective action
- Hydrodynamic effective action
- thermo-field dynamics
- Eternal AdS black hole

1. Closed system of IR modes

Closed: conserved energy of the IR modes

1) harmonic oscillator: $\ddot{X}(t) + \omega^2 X(t) = 0$ $S = \int dt \left(\frac{1}{2} m \dot{X}^2 - \frac{1}{2} m \omega^2 X^2 \right)$

2) ideal fluid $\partial_\mu s^\mu = 0$


[Dubovsky, Gregoire, Nicolis, and Rattazzi JHEP 0512260]

$$S_0 = \int d^4x F(b), \quad b \equiv \sqrt{\det \partial_\mu \phi^I \partial^\mu \phi^J}$$
$$J^\mu \equiv \frac{1}{6} \epsilon^{\mu\alpha\beta\gamma} \epsilon_{IJK} \partial_\alpha \phi^I \partial_\beta \phi^J \partial_\gamma \phi^K$$
$$b^2 = -J_\mu J^\mu, \quad u^\mu = \frac{1}{b} J^\mu$$

2. Open system of IR modes: dissipation

Open system:

energy flow


$$\{\psi\} = \{\phi\}_{\text{IR}} + \{\chi\}_{\text{UV}}$$

1) damped harmonic oscillator

$$\ddot{X}(t) + \gamma\dot{X}(t) + \omega^2 X(t) = 0$$

?

2) viscous fluid

$$\text{effective action } \partial_\mu s^\mu > 0$$

dissipative fluid

?

3. Non-equilibrium dynamics

Suppose the system at $t_i \rightarrow -\infty$ density

$$\rho(t_i)$$

If the system evolves

with a time independent. : H

$$\rho(t) = e^{-iH(t-t_i)} \rho(t_i) e^{iH(t-t_i)}$$

If H time-dependent:

$$H(t)$$

$$\rho(t) = U(t, t_i) \rho(t_i) U(t_i, t)$$

$$i \frac{\partial U(t, t')}{\partial t} = H(t) U(t, t')$$

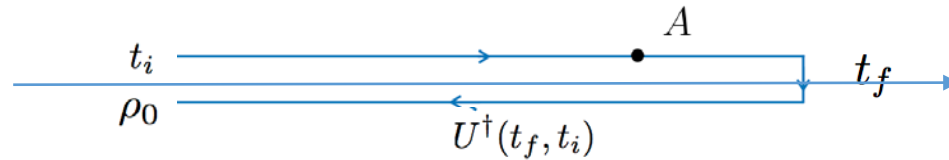
$$U(t, t') = T \left(e^{-i \int_{t'}^t dt'' H(t'')} \right)$$

4. Schwinger-Keldysh contour (CTP)

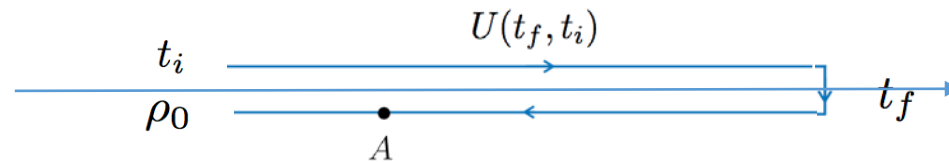
Time evolution of operators

$$\begin{aligned}\langle A \rangle(t) &= \text{Tr}(\rho(t)A) \\ &= \text{Tr} \left(U(t, t_i) \rho(t_i) U(t_i, t) A \right) \\ &= \text{Tr} \left(\rho(t_i) U(t_i, t) A U(t, t_i) \right) \\ &= \text{Tr} \left(\rho(t_i) U(t_i, t) A U(t, t_f) U(t_f, t_i) \right)\end{aligned}$$

$t_f \rightarrow +\infty$



$$= \text{Tr} \left(\rho(t_i) U(t_i, t_f) U(t_f, t) A U(t, t_i) \right)$$

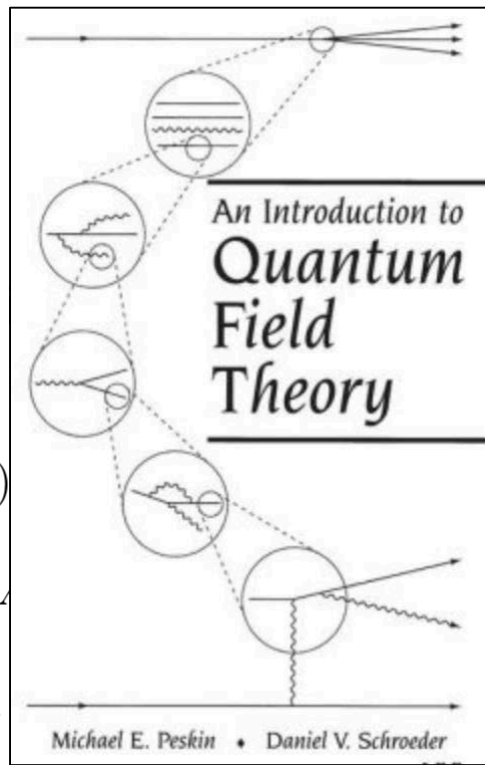


Closed time path

5. Comparison with ?

Let us suppose the system initially

$$\begin{aligned} \langle A \rangle(t) &= \text{Tr}(\rho(t)A) \\ &= \text{Tr} \left(U(t, t_i) \rho(t_i) U(t_i, t) \right) \\ &= \sum_n \langle n | 0_{t_i} \rangle \langle 0_{t_i} | U(t_i, t) A U(t, t_f) | n \rangle \\ &= \langle 0_{t_i} | U(t_i, t_f) U(t_f, t) A U(t, t_i) | 0_{t_i} \rangle \end{aligned}$$



$$\rho(t_i) = |0_{t_i}\rangle\langle 0_{t_i}|$$

$U(t, t_i)$

Adiabatic assumption

$$U(t_f, t_i) |0_{t_i}\rangle = e^{iL} |0_{t_f}\rangle \quad \rightarrow \quad \langle 0_{t_f} | U(t_f, t_i) | 0_{t_i} \rangle = e^{iL}$$

$$= e^{-iL} \langle 0_{t_f} | U(t_f, t) A U(t, t_i) | 0_{t_i} \rangle$$

$$= \frac{\langle 0_{t_f} | U(t_f, t) A U(t, t_i) | 0_{t_i} \rangle}{\langle 0_{t_f} | U(t_f, t_i) | 0_{t_i} \rangle}$$



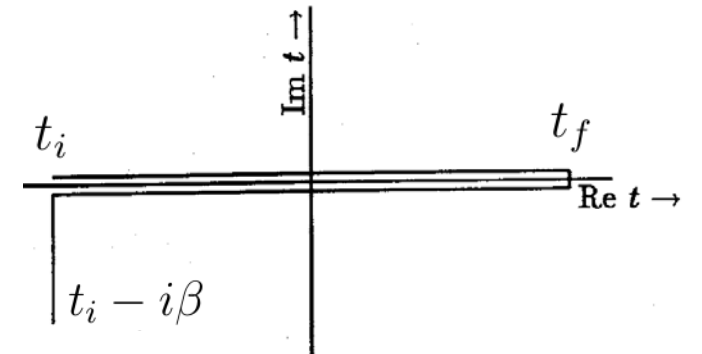
$$\langle \Omega | T \{ \phi(x) \phi(y) \} | \Omega \rangle = \lim_{T \rightarrow \infty (1-i\epsilon)} \frac{\langle 0 | T \left\{ \phi_I(x) \phi_I(y) \exp \left[-i \int_{-T}^T dt H_I(t) \right] \right\} | 0 \rangle}{\langle 0 | T \left\{ \exp \left[-i \int_{-T}^T dt H_I(t) \right] \right\} | 0 \rangle} \quad (4.31)$$

6. Dynamics around thermal equilibrium

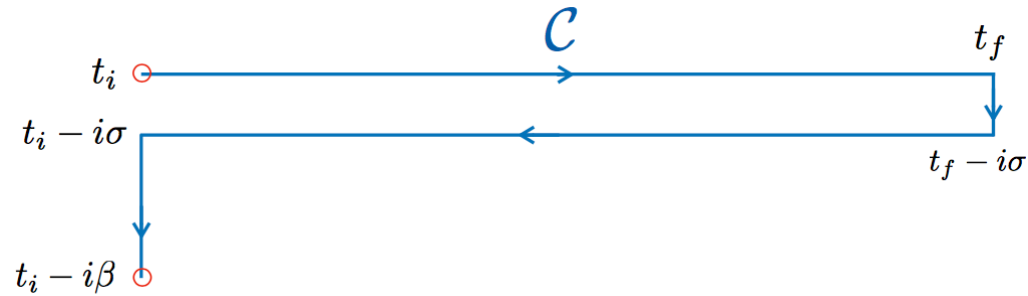
Let us suppose the system initially in thermal equilibrium:

$$\rho(t_i) = \frac{e^{-\beta H_i}}{\text{Tr } e^{-\beta H_i}} = \frac{U(T - i\beta, T)}{\text{Tr } U(T - i\beta, T)}$$
$$H(t) = \begin{cases} H_i & \text{for } \text{Re } t \leq 0 \\ \mathcal{H}(t) & \text{for } \text{Re } t \geq 0 \end{cases}$$

$$\langle A \rangle(t) = \text{Tr}(\rho(t)A)$$
$$= \frac{\text{Tr} (U(t_i - i\beta, t_i) U(t_i, t_f) U(t_f, t) A U(t, t_i))}{\text{Tr} (U(t_i - i\beta, t_i) U(t_i, t_f) U(t_f, t_i))}$$



7. General SK contour around equilibrium




$$\sigma \in [0, \beta)$$

While the correlation functions are different compared to those of the $\sigma = 0$ case, they are simply related to each other by analytic continuations

8. Generating functional

Rev of generating functional in zero temperature QFT

$$Z[J] = e^{iW[J]} = \int \mathcal{D}\phi e^{i \int d^4x (\mathcal{L} + J\phi)}$$

$$\langle \phi(x)\phi(y) \rangle_F \sim \frac{\delta^2 \ln Z[J]}{\delta J(x)\delta J(y)}$$

On the other hand

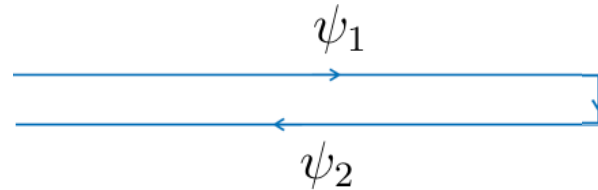
$$\langle \mathcal{O} \rangle = \int \mathcal{D}\phi e^{i \int d^4x \mathcal{L}} \mathcal{O} = \text{Tr} (\rho(t_i) \mathcal{O})$$

$$Z[J] = e^{iW[J]} = \int \mathcal{D}\phi e^{i \int d^4x \mathcal{L}} e^{i \int d^4x J\phi} = \langle \mathcal{T} e^{i \int d^4x J\phi} \rangle = \text{Tr} (\rho(t_i) \mathcal{T} e^{i \int d^4x J\phi})$$

$$Z[J] = e^{iW[J]} = \text{Tr} (\rho(t_i) \mathcal{T} e^{i \int d^4x J\phi})$$

9. SK formalism: generating functional and effective action on CTP

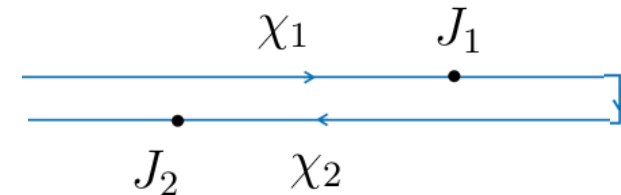
Doubling the dynamical fields at the micro level:



$$\begin{aligned} \langle \cdots \rangle &= \text{Tr} (\rho_0 \cdots) = \int_{\rho_0} \mathcal{D}\psi_1 \mathcal{D}\psi_2 e^{iS[\psi_1] - iS[\psi_2] + i \ln \cdots} \\ &= \int \mathcal{D}\chi_1 \mathcal{D}\chi_2 e^{iS_{\text{EF}}[\chi, \chi_2; \rho_0] + i \ln \cdots} \end{aligned}$$

Connected correlation functions defined on a CTP contour:

$$\begin{aligned} Z[J] &= e^{iW[J]} = \text{Tr} \left(\rho_0 \mathcal{P} e^{i \int d^4x (J_1(x)\chi_1(x) - J_2(x)\chi_2(x))} \right) \\ &= \text{Tr} (U_1(+\infty, -\infty; J_1) \rho_0 U_2(-\infty, +\infty; J_2)) \end{aligned}$$



10. r-a basis

It is more convenient to work with

$$\begin{aligned}\phi_r &= \frac{\phi_1 + \phi_2}{2} & \phi_a &= \phi_1 - \phi_2 \\ J_r &= \frac{J_1 + J_2}{2} & J_a &= J_1 - J_2\end{aligned}$$

$$\begin{aligned}e^{iW[J]} &= \text{Tr} \left(\rho_0 \mathcal{P} e^{i \int d^4x (J_1(x)\chi_1(x) - J_2(x)\chi_2(x))} \right) \\ &= \text{Tr} \left(\rho_0 \mathcal{P} e^{i \int d^4x (J_a(x)\chi_r(x) + J_r(x)\chi_a(x))} \right)\end{aligned}$$

$$\langle \chi_i(x)\chi_j(y) \rangle \sim \frac{\delta^2 W[J_1, J_2]}{\delta J_i(x)\delta J_j(y)}, \quad i, j \in \{1, 2\} \quad \langle \chi_{\alpha_1}(x)\chi_{\alpha_2}(y) \rangle \sim \frac{\delta^2 W[J_r, J_a]}{\delta J_{\bar{\alpha}_1}(x)\delta J_{\bar{\alpha}_2}(y)}, \quad \alpha_1, \alpha_2 \in \{r, a\}$$

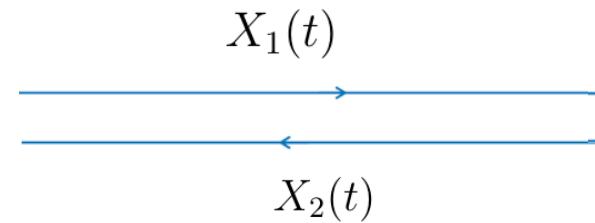
$$\begin{aligned}W[J_1, J_2] &= \frac{1}{2} \int d^4x d^4y \begin{pmatrix} J_1(x) & J_2(x) \end{pmatrix} \begin{pmatrix} G_F & -iG_- \\ -iG_+ & G_F \end{pmatrix} \begin{pmatrix} J_1(y) \\ J_2(y) \end{pmatrix} \\ &= \frac{1}{2} \int d^4x d^4y \begin{pmatrix} J_r(x) & J_a(x) \end{pmatrix} \begin{pmatrix} 0 & G_A \\ G_R & iG_S \end{pmatrix} \begin{pmatrix} J_r(y) \\ J_a(y) \end{pmatrix}\end{aligned}$$

11. Dissipation from EFT

Symmetries of the generating functional and effective action:

$$S_{\text{EF}; \text{CTP}}[\chi_1, \chi_2] = -S_{\text{EF}; \text{CTP}}^*[\chi_2, \chi_1]$$

e.g. damped Harmonic oscillator:



$$S_{\text{EF}; \text{CTP}}[X_1, X_2] = S[X_1] - S[X_2] + S_{\text{int}}[X_1, X_2]$$

$$S[X] = \int dt \left(\frac{1}{2} m \dot{X}^2 - \frac{1}{2} m \omega^2 X^2 \right) \quad S_{\text{int}}[X_1, X_2] = \int dt \frac{1}{2} \gamma \left(\dot{X}_1 X_2 - X_1 \dot{X}_2 \right)$$

$$\ddot{X}_{1,2} + \gamma \dot{X}_{2,1} + \omega^2 X_{1,2} = 0$$

$$X_1(t_f) = X_2(t_f)$$



$$X_1(t) = X_2(t) = X(t)$$

$$\ddot{X} + \gamma \dot{X} + \omega^2 X = 0$$

12. Effective action for dissipative fluids

Symmetries:

[1511.03646, JHEP, Crossley, Glorioso and Liu]

- 1) KMS
- 2) Microscopic time reversal
- 3) Unitarity
- 4) Fluid world volume diff

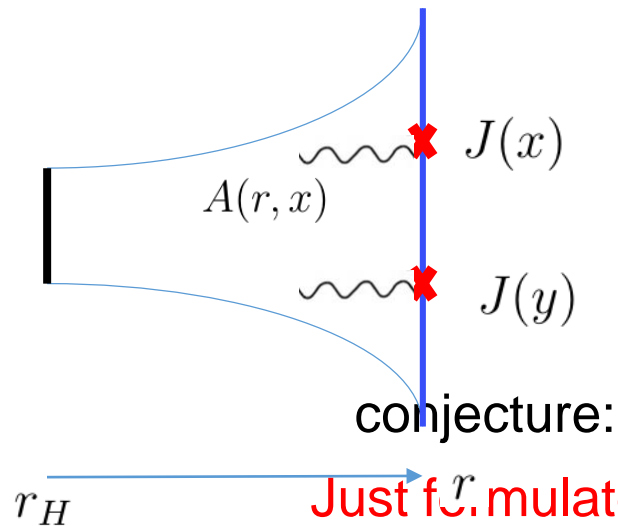
Fixe the form of the effective action

The EoMs read: $\partial_\mu T^{\mu\nu} = 0$

$$T^{\mu\nu} = (\epsilon + p)u^\mu u^\nu + p\eta^{\mu\nu} - \eta P^{\mu\alpha} P^{\nu\beta} (\partial_\alpha u_\beta + \partial_\beta u_\alpha) - \left(\zeta - \frac{2}{3}\eta\right) P^{\mu\nu} \partial \cdot u$$

12+1. Dual gravity picture

Boundary thermal correlation functions from gravity



$$S = -\frac{1}{16\pi G_5} \int_{\mathcal{M}} d^5x \sqrt{-g} \left(R - \frac{12}{L^2} + F^{MN} F_{MN} \right)$$



$$\nabla_{\mu} F^{\mu\nu} = 0 \quad \rightarrow \quad A_{sol}^{\mu}$$

$$Z_{grav.}[A_{sol}^{\mu}] = e^{iS[A_{sol}^{\mu}]} = Z_{CFT}[J]$$

[Witten, 9802150]

Making analytic continuation to the Minkowski needs to knowing the A_{sol}^{μ} all Matsubara frequencies.

Practically impossible

14. Son-Starinets prescription [0205051, JHEP]

The on-shell action is formally written as

$$S[A_{sol}^\mu] = \int \frac{d^4 p}{(2\pi)^4} A_0^\mu(-p) \mathcal{G}_{\mu\nu}(p, r) A_0^\nu(p) \Big|_{r=r_H}^{r \rightarrow \infty}$$

Its derivative with respect to the boundary value of A_{sol}^μ reads

$$G(p) = \mathcal{G}_{\mu\nu}(p, r) \Big|_{r=r_H}^{r \rightarrow \infty} - \mathcal{G}_{\mu\nu}(-p, r) \Big|_{r=r_H}^{r \rightarrow \infty}$$

In contrast to this G_R quantity is real!

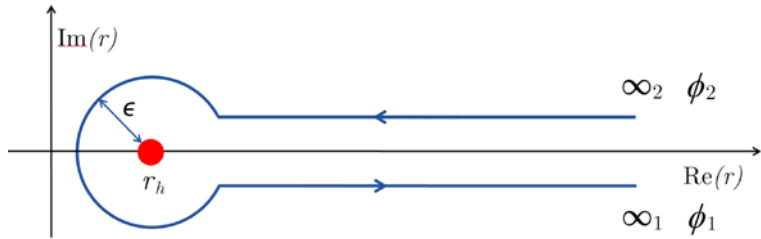
Recipe: $G_R(p) = -2 \lim_{r \rightarrow \infty} \mathcal{G}_{\mu\nu}(p, r)$ [Son, Starinets]
[0205052, JHEP, Policastro, Son, Starinets]

Towards the dual description of the SK contour in gravity

[0805.0150, PRL, Skenderis, Van Rees], [1504.07616, JHEP, de Boer, Heller, Pinzani-Fokeeva]

15. Very recent prescription:

Effective action. [1812.08785, Glorioso, Crossley, Liu]



$$ds^2 = -f(r, x^\mu)dv^2 + 2drdv + \lambda_{ij}(r, x^\mu)dx^i dx^j$$

$$f(r \rightarrow \infty) \rightarrow r^2, \quad \lambda_{ij}(r \rightarrow \infty) \rightarrow \delta_{ij}r^2$$

$$S = -\frac{1}{2} \int d^{d+1}x \sqrt{-g} g^{MN} \partial_M \Phi \partial_N \Phi$$

$$\frac{1}{\sqrt{\lambda}} \partial_r (f \sqrt{\lambda} \partial_r \Phi) + \partial_0 \partial_r \Phi + \frac{1}{\sqrt{\lambda}} \partial_r (\sqrt{\lambda} \partial_0 \Phi) + \lambda^{ij} \partial_i \partial_j \Phi = 0$$

$$\Phi(r, x^\mu) \rightarrow \phi_1(x), \quad r \rightarrow \infty_1$$

$$\Phi(r, x^\mu) \rightarrow \phi_2(x), \quad r \rightarrow \infty_2$$

Derivative expansion

$$\Phi = \Phi^{(0)} + \Phi^{(1)} + \Phi^{(2)} + \dots$$

$$\phi_r = \frac{1}{2}(\phi_1 + \phi_2)$$

$$\phi_a = \phi_1 - \phi_2$$

16. Very recent prescription:

The on-shell action to 2nd order in boundary derivatives then reads

$$\begin{aligned} W &= W^{(0)} + W^{(1)} + W^{(2)} \\ &= \frac{i}{\beta} \int d^d x \sqrt{\lambda_h} \phi_a^2 - \int d^d x \sqrt{\lambda_h} \phi_a \partial_0 \phi_r \\ &\quad + \int d^d x \left\{ Q_{ra} \partial_0 \phi_a \partial_0 \phi_r + Q_{ra}^{ij} \partial_i \phi_a \partial_j \phi_r + \frac{i}{2\beta} Q_{aa} (\partial_0 \phi_a)^2 + \frac{i}{2\beta} Q_{aa}^{ij} \partial_i \phi_a \partial_j \phi_a \right\} \end{aligned}$$

$$Q_{ra} = \frac{1}{2} r_h^2 (\log 2 - 1), \quad Q_{ra}^{ij} = \frac{1}{2} r_h^2 \delta_{ij}$$

$$= \frac{1}{2} \int d^4 x d^4 y \begin{pmatrix} J_r(x) & J_a(x) \end{pmatrix} \begin{pmatrix} 0 & G_A \\ G_R & iG_S \end{pmatrix} \begin{pmatrix} J_r(y) \\ J_a(y) \end{pmatrix}$$

$$G_R(\omega, q) = r_h^3 i\omega + \frac{1}{2} r_h^2 (\log 2 - 1) \omega^2 + \frac{1}{2} r_h^2 q^2$$

17. Is the thermal state like a vacuum state?

Average in thermal state: $\langle A \rangle_\beta = Z^{-1}(\beta) \text{Tr} e^{-\beta \mathcal{H}} A$

Could it be also written as $\langle A \rangle_\beta = \langle 0, \beta | A | 0, \beta \rangle$ vacuum state $|0, \beta\rangle$

If yes, then

$$|0, \beta\rangle = \sum_n |n\rangle \langle n | 0, \beta\rangle = \sum_n f_n(\beta) |n\rangle$$

so

$$\langle 0, \beta | A | 0, \beta \rangle = \sum_{n,m} f_n^*(\beta) f_m(\beta) \langle n | A | m \rangle$$



$$f_n^*(\beta) f_m(\beta) = Z^{-1}(\beta) e^{-\beta E_n} \delta_{nm}$$

and since

$$\langle A \rangle_\beta = Z^{-1}(\beta) \sum_n e^{-\beta E_n} \langle n | A | n \rangle$$



With restricting our self to the original Hilbert space, we cannot define a thermal vacuum which gives

$$\langle A \rangle_\beta = Z^{-1}(\beta) \sum_n e^{-\beta E_n} \langle n | A | n \rangle$$

18. Thermo-field dynamics: thermal vacuum

Let us assume the product space of two systems

$$|n, \tilde{m}\rangle = |n\rangle \otimes |\tilde{m}\rangle$$

Now expand the thermal state as

$$|0, \beta\rangle = \sum_n f_n(\beta) |n, \tilde{n}\rangle = \sum_n f_n(\beta) |n\rangle \otimes |\tilde{n}\rangle$$

$$\begin{aligned} \langle 0, \beta | A | 0, \beta \rangle &= \sum_{n, m} f_n^*(\beta) f_m(\beta) \langle n, \tilde{n} | A | m, \tilde{m} \rangle \\ &= \sum_{n, m} f_n^*(\beta) f_n(\beta) \langle n | A | n \rangle \delta_{nm} \\ &= \sum_n f_n^*(\beta) f_n(\beta) \langle n | A | n \rangle \end{aligned}$$

Compare it with

$$\langle A \rangle_\beta = Z^{-1}(\beta) \sum_n e^{-\beta E_n} \langle n | A | n \rangle$$

thermal vacuum state

$$f_n^*(\beta) f_n(\beta) = Z^{-1}(\beta) e^{-\beta E_n}$$

$$|0, \beta\rangle = \frac{1}{\sqrt{Z(\beta)}} \sum_n e^{-\beta E_n / 2} |n, \tilde{n}\rangle$$

19. Thermo-field dynamics: harmonic oscillator

Let us consider a fermionic harmonic oscillator:

$$H = \omega a^\dagger a \quad [a, a^\dagger]_+ = 1 \quad |1\rangle = a^\dagger |0\rangle$$

Make a copy:

$$\tilde{H} = \omega \tilde{a}^\dagger \tilde{a}$$

Space of states

$$|0, \tilde{0}\rangle; |0, \tilde{1}\rangle; |1, \tilde{0}\rangle; |1, \tilde{1}\rangle$$

Thermal vacuum

$$|0, \beta\rangle = \frac{1}{\sqrt{1 + e^{-\beta\omega}}} (|0, \tilde{0}\rangle + e^{-\beta\omega/2} |1, \tilde{1}\rangle)$$

Thermal state can be constructed out of state: $|0, \tilde{0}\rangle$

$$U(\theta) |0, \tilde{0}\rangle = \cos \theta(\beta) |0, \tilde{0}\rangle + \sin \theta(\beta) |1, \tilde{1}\rangle = |0, \beta\rangle$$

$$U(\theta) = e^{-\theta(\beta)(\tilde{a}a - a^\dagger \tilde{a}^\dagger)}$$

We find the transformation of the ladder operators:

$$a(\beta) = a \cos \theta(\beta) - \tilde{a}^\dagger \sin \theta(\beta) \quad \text{um state } a(\beta) |0, \beta\rangle = 0$$

$$\tilde{a}(\beta) = \tilde{a} \cos \theta(\beta) + a^\dagger \sin \theta(\beta)$$

20. Thermo-field dynamics: Hamiltonian

Thermal Hilbert space of oscillator

$$|0, \beta\rangle; a^\dagger(\beta)|0, \beta\rangle; \bar{a}^\dagger(\beta)|0, \beta\rangle; a^\dagger(\beta)\bar{a}^\dagger(\beta)|0, \beta\rangle$$

These states are not eigen states of H or \tilde{H}

But since

$$a^\dagger(\beta)a(\beta) - \bar{a}^\dagger(\beta)\bar{a}(\beta) = a^\dagger a - \bar{a}^\dagger \bar{a}$$

So the Hamiltonian of the combined system is

$$\hat{H} = H - \tilde{H}$$

Is there any relation between “thermo-field” and “CTP”?

21. Thermo-field dynamics: free KG theory

Let's consider

$$\begin{aligned}\mathcal{L} &= \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{m^2}{2} \phi^2 \\ \tilde{\mathcal{L}} &= \frac{1}{2} \partial_\mu \tilde{\phi} \partial^\mu \tilde{\phi} - \frac{m^2}{2} \tilde{\phi}^2\end{aligned}\quad \hat{\mathcal{L}} = \mathcal{L} - \tilde{\mathcal{L}}$$

Green's function is given by a 2×2 matrix

$$iG(x-y) = \langle 0, \tilde{0} | T(\Phi(x)\Phi(y)) | 0, \tilde{0} \rangle \quad G(k) = \begin{pmatrix} \frac{1}{k^2 - m^2 + i\epsilon} & 0 \\ 0 & -\frac{1}{k^2 - m^2 - i\epsilon} \end{pmatrix}$$

The Green's function in thermal vacuum

$$\begin{aligned}iG_\beta(x-y) &= \langle 0, \beta | T(\Phi(x)\Phi(y)) | 0, \beta \rangle \\ &= \langle 0, \tilde{0} | T(U^\dagger(\theta)\Phi(x)\Phi(y)U(\theta)) | 0, \tilde{0} \rangle \\ &= \bar{U}(-\theta) iG(x-y) \bar{U}^T(-\theta)\end{aligned}$$

$$G_\beta(k) = \begin{pmatrix} \frac{1}{k^2 - m^2 + i\epsilon} & 0 \\ 0 & -\frac{1}{k^2 - m^2 - i\epsilon} \end{pmatrix} - 2i\pi n_B(|k^0|) \delta(k^2 - m^2) \begin{pmatrix} 1 & e^{\beta|k^0|/2} \\ e^{\beta|k^0|/2} & 1 \end{pmatrix}$$

22. thermo-field vs. CPT

1. Propagator of KG field on CPT

$$G = \begin{pmatrix} G_{++} & G_{+-} \\ G_{-+} & G_{--} \end{pmatrix} \begin{matrix} G_{++}(p) = \left(\frac{1}{p^2 - m^2 + i\epsilon} - 2i\pi n_B(|p^0|) \delta(p^2 - m^2) \right) \\ G_{+-}(p) = -2i\pi \left(\theta(-p^0) + n_B(|p^0|) \right) \delta(p^2 - m^2) \end{matrix}$$

2. Propagator of KG field on Generalized CTP

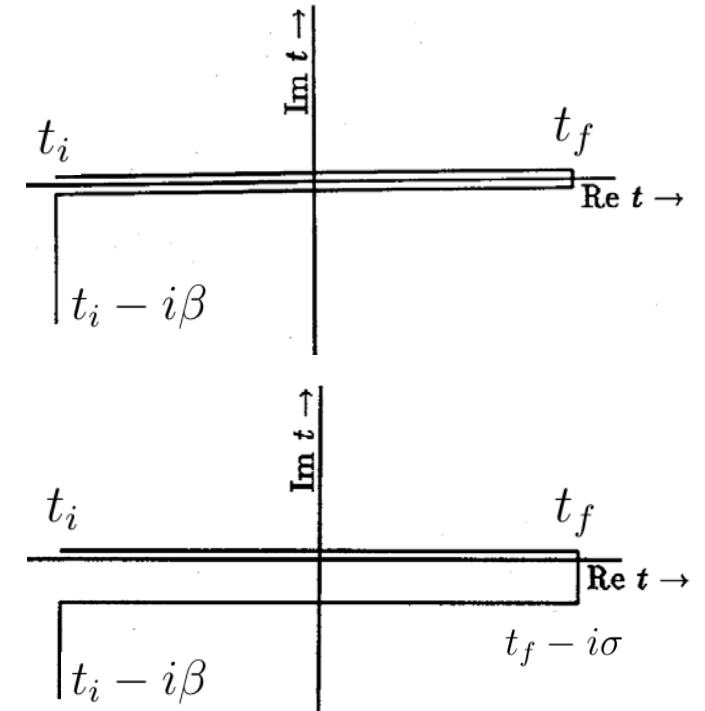
$$\begin{aligned} G_{11}(p) &= G_{++}(p) \\ G_{12}(p) &= e^{\sigma p^0} G_{+-}(p) \end{aligned}$$

3. Propagator of KG field on Generalized CTP

$$G_{\beta}(k) = \begin{pmatrix} \frac{1}{k^2 - m^2 + i\epsilon} & 0 \\ 0 & -\frac{1}{k^2 - m^2 - i\epsilon} \end{pmatrix} - 2i\pi n_B(|k^0|) \delta(k^2 - m^2) \begin{pmatrix} 1 & e^{\beta|k^0|/2} \\ e^{\beta|k^0|/2} & 1 \end{pmatrix}$$

Thermo-field dynamics = CTP formalism with

$$\sigma = \frac{\beta}{2}.$$



23. Analogy with eternal AdS black hole I

Let's consider a "Schwartz-child" black-hole

$$ds^2 = -f(r)dt^2 + \frac{1}{f(r)}dr^2 + r^2 d\vec{x}^2$$

$$f(r) = r^2 \left(1 - \frac{r_H^4}{r^4}\right)$$

Kruskal coordinates:

$$ds^2 = d\vec{x}^2 = 0 \quad \rightarrow \quad dt = \frac{1}{f(r)} dr \equiv dr^*$$

$$v = t + r_*$$

$$u = t - r_*$$

R:

$$U = -e^{-\frac{2\pi}{\beta} u}$$

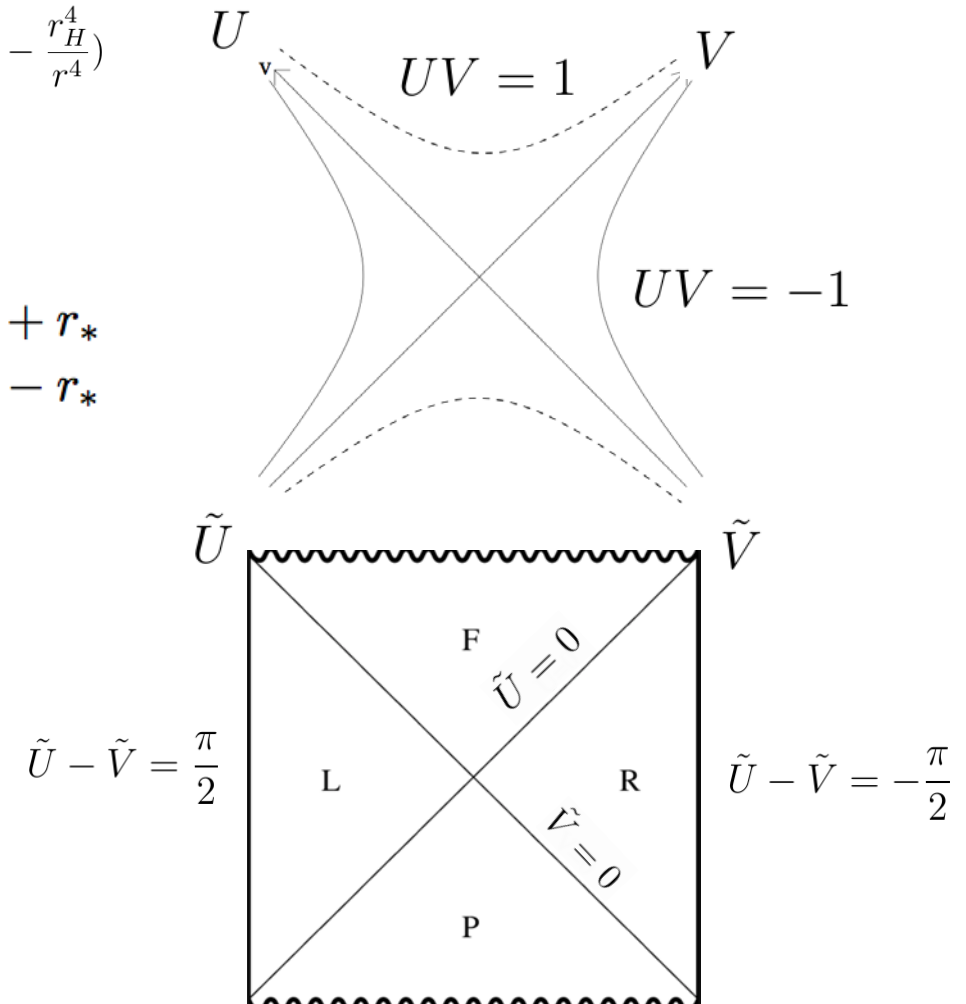
$$V = e^{\frac{2\pi}{\beta} v}$$

Penrose diagram:

R:

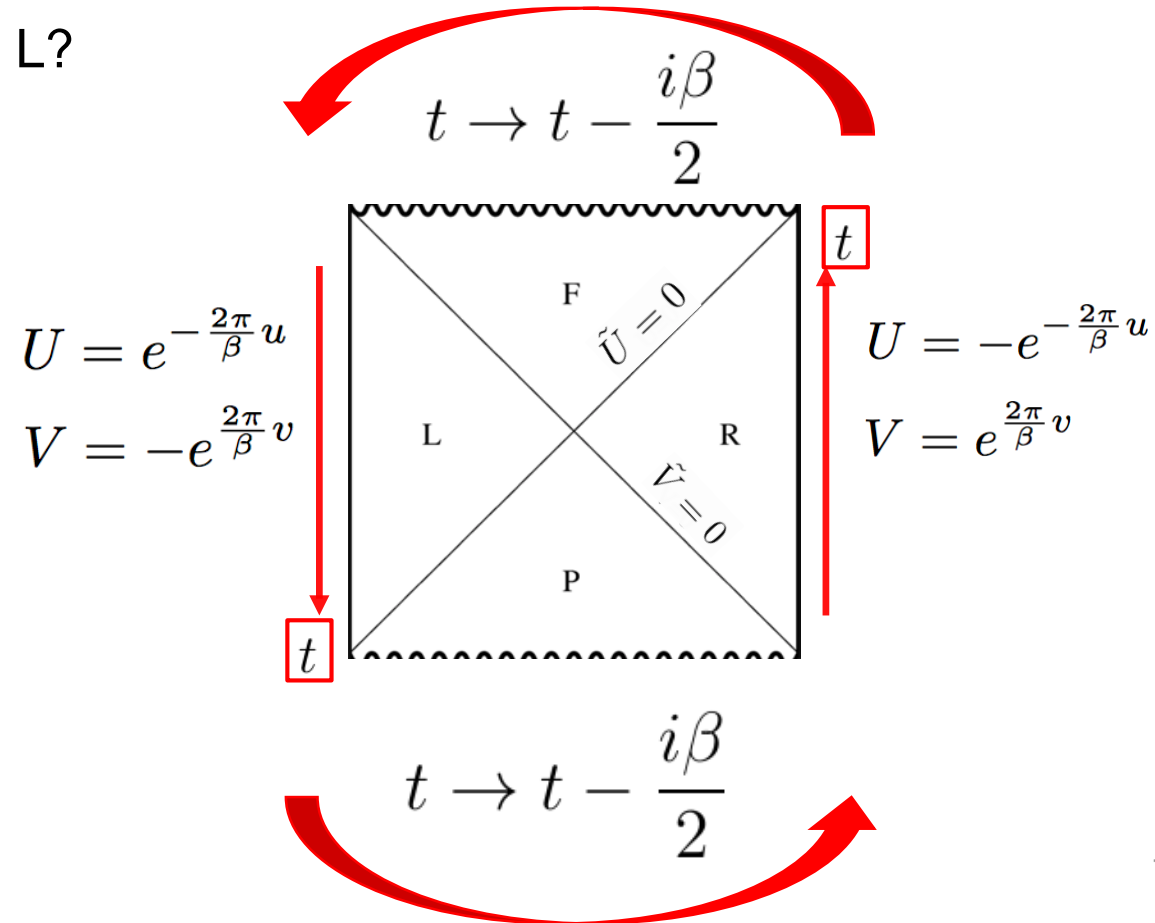
$$\tilde{U} = \tan^{-1} U$$

$$\tilde{V} = \tan^{-1} V$$



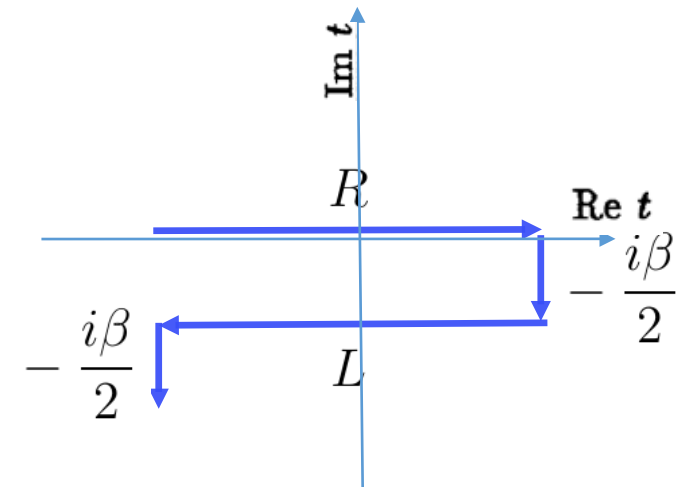
24. Analogy with eternal AdS black hole II

How to go from R to L?



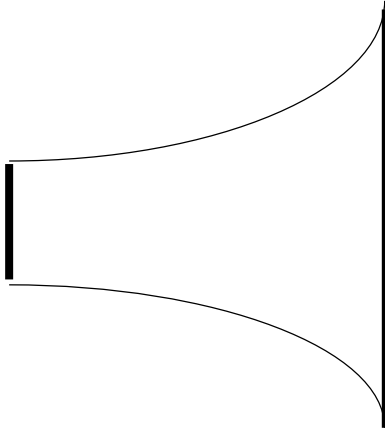
$$v = t + r_*$$

$$u = t - r_*$$



25. Summary

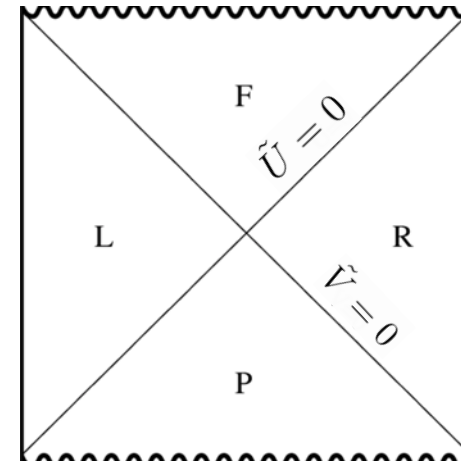
One-sided AdS black hole



Mixed thermal state

$$\rho = \text{Tr} e^{-\beta H}$$

eternal black-hole



thermo-field state

$$|0, \beta\rangle$$

Thank you for your attention