

# Hidden Conformal Symmetry of Warped $\text{AdS}_3$ Black Holes

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# Warped AdS<sub>3</sub> spacetime

- Warped AdS<sub>3</sub> spacetimes appear in the near horizon geometry of extremal Kerr black holes:
- The Kerr metric is

$$ds^2 = -\frac{\Delta}{\rho^2} \left( dt - a \sin^2 \theta d\phi \right)^2 + \frac{\rho^2}{\Delta} d\hat{r}^2 + \rho^2 d\theta^2 + \frac{\sin^2 \theta}{\rho^2} \left( (\hat{r}^2 + a^2) d\hat{\phi} - a dt \right)^2 \quad (1)$$

$$a \equiv \frac{J}{M}, \quad \Delta \equiv \hat{r}^2 - 2Mr + a^2, \quad \rho^2 \equiv \hat{r}^2 + a^2 \cos^2 \theta, \quad (2)$$

- Horizons are at

$$r_{\pm} = M \pm \sqrt{M^2 - a^2} \quad (3)$$

- For extremal cases (i.e.  $r_+ = r_-$ ), we define new coordinates

$$t = \frac{\lambda \hat{t}}{2M^2}, \quad r = \frac{\hat{r} - M}{\lambda}, \quad \phi = \hat{\phi} - \frac{\hat{t}}{2M}. \quad (4)$$

- In the  $\lambda \rightarrow 0$  limit, the result is the near-horizon geometry of the extremal Kerr (NHEK):

$$ds^2 = \Gamma(\theta) \left[ -r^2 dt^2 + \frac{dr^2}{r^2} + d\theta^2 \right] + \gamma(\theta) (d\phi + r dt)^2, \quad (5)$$

where

$$\Gamma(\theta) = M^2 (1 + \cos^2 \theta), \quad \gamma(\theta) = \frac{4M^2 \sin^2 \theta}{1 + \cos^2(\theta)}. \quad (6)$$

- The NHEK geometry is not asymptotically flat!
- This geometry has  $U(1) \times SL(2, R)$  isometry. The rotational  $U(1)$  isometry is generated by the Killing vector  $J_0 = -\partial_\varphi$  and the  $SL(2, R)$  is generated by

$$\begin{aligned}
 \bar{J}_1 &= \partial_t , \\
 \bar{J}_2 &= t\partial_t - r\partial_r , \\
 \bar{J}_3 &= \left( \frac{1}{2r^2} + \frac{t^2}{2} \right) \partial_t - tr\partial_r - \frac{k}{r}\partial_\phi .
 \end{aligned} \tag{7}$$

- For fixed  $\theta$ , the geometry (5) is similar to the three-dimensional warped AdS<sub>3</sub> spacetime,

$$ds^2 = \frac{\ell^2}{(\nu^2 + 3)} \left[ -\cosh^2 \sigma d\tau^2 + d\sigma^2 + \frac{4\nu^2}{\nu^2 + 3} (du + \sinh \sigma d\tau)^2 \right] \quad (8)$$

which is a solution of the Topologically Massive Gravity (TMG):

$$I_{TMG} = \frac{1}{16\pi G} \int_{\mathcal{M}} d^3x \sqrt{-g} (R + 2/\ell^2) + \frac{\ell}{96\pi G\nu} \int_{\mathcal{M}} d^3x \sqrt{-g} \varepsilon^{\lambda\mu\nu} \Gamma_{\lambda\sigma}^r \left( \partial_\mu \Gamma_{r\nu}^\sigma + \frac{2}{3} \Gamma_{\mu\tau}^\sigma \Gamma_{\nu r}^\tau \right) \quad (9)$$

- Besides WAdS<sub>3</sub> geometries, the action (9) has warped AdS<sub>3</sub> black hole solutions. These black holes are discrete quotients of warped AdS<sub>3</sub> just as BTZ black holes are discrete quotients of the ordinary AdS<sub>3</sub>.
- The metric of WAdS<sub>3</sub> black hole is

$$\begin{aligned}
 ds^2 &= -N(r)^2 dt^2 + \ell^2 R(r)^2 (d\phi + N^\phi(r) dt)^2 + \frac{\ell^4 dr^2}{4R(r)^2 N(r)^2}, \\
 N(r)^2 &\equiv \frac{\ell^2 (\nu^2 + 3)(r - r_+)(r - r_-)}{4R(r)^2}, \\
 N^\phi(r) &\equiv \frac{2\nu r - \sqrt{r_+ r_-} (\nu^2 + 3)}{2R(r)^2}, \\
 R(r)^2 &\equiv \frac{r}{4} \left( 3(\nu^2 - 1)r + (\nu^2 + 3)(r_+ + r_-) \right. \\
 &\quad \left. - 4\nu \sqrt{r_+ r_-} (\nu^2 + 3) \right). \tag{10}
 \end{aligned}$$

# WAdS<sub>3</sub>/CFT<sub>2</sub> conjecture

- Quantum TMG is holographically dual to a 2D boundary CFT with  $c_R = \frac{(5\nu^2+3)\ell}{G\nu(\nu^2+3)}$  and  $c_L = \frac{4\nu\ell}{G(\nu^2+3)}$ .  
[Anninos, Li, Padi, Song and Strominger (2008)]
- The motivations for this conjecture are that application of the Cardy formula to the CFT density of states reproduces the black hole entropy.
- The close relation between WAdS<sub>3</sub> and NHEK geometries results in the following conjecture which is known as "Kerr/CFT correspondence":

Quantum gravity in the background of near horizon extremal Kerr black holes is holographically dual to a chiral two-dimensional conformal field theory.

[Guica, Hartman, Song, Strominger (2008)]



# Hidden Conformal Symmetry

- The isometry of  $WAdS_3$  and NHEK is  $U(1) \times SL(2, R)$ . They do not have the  $SL(2, R) \times SL(2, R)$  conformal symmetry of the dual CFT.
- The attempts to enhance the  $U(1) \times SL(2, R)$  isometry of  $WAdS_3$  to the full conformal symmetry by making use of the Brown and Henneaux's asymptotic symmetry method, have not been unambiguously successful.
- The question which still remains unanswered is that What is the interpretation of dual conformal symmetry in terms of gravitational bulk theory?

- "Hidden conformal symmetry" may provide an answer to this question:
- In this proposal the conformal symmetry is not derived from a conformal symmetry of the spacetime geometry. It is a symmetry of solution space:
- To see this, Consider the warped AdS<sub>3</sub> black hole (10). We want to study the propagation of a massive scalar field in this background.
- The Klein-Gordon equation for a massive scalar with mass  $m$  is

$$\left( \frac{1}{\sqrt{-g}} \partial_\mu \sqrt{-g} \partial^\mu - m^2 \right) \Phi = 0. \quad (11)$$

- Expanding in eigenmodes

$$\Phi(t, r, \theta, \phi) = e^{-i\omega t + ik\phi} \Phi(r, \theta), \quad (12)$$

and imposing

$$\omega^2 \ll \frac{(\nu^2 + 3)^2}{3(\nu^2 - 1)}. \quad (13)$$

simplifies the wave equation (11) to the form

$$\tilde{\mathcal{H}}^2 \Phi = \mathcal{H}^2 \Phi = \frac{\ell^2 m^2}{\nu^2 + 3} \Phi \quad (14)$$

where  $\tilde{\mathcal{H}}^2$  and  $\mathcal{H}^2$  are the quadratic Casimir of two following  $SL(2, R)$ s: **[R.F. (2010)]**

$$H_0 = -\frac{2i\nu}{\nu^2 + 3} \frac{T_L}{T_R} \partial t + \frac{i}{2\pi\ell T_R} \partial_\phi,$$

$$H_1 = i e^{-2\pi\ell T_R \phi} \left[ -\frac{\nu}{(\nu^2 + 3)\sqrt{\Delta}} \left( (2r - r_+ - r_-) \frac{T_L}{T_R} + r_+ - r_- \right) \partial_t \right. \\ \left. + \sqrt{\Delta} \partial_r + \frac{2r - r_+ - r_-}{4\pi\ell T_R \sqrt{\Delta}} \partial_\phi \right],$$

$$H_{-1} = i e^{2\pi\ell T_R \phi} \left[ -\frac{\nu}{(\nu^2 + 3)\sqrt{\Delta}} \left( (2r - r_+ - r_-) \frac{T_L}{T_R} + r_+ - r_- \right) \partial_t \right. \\ \left. - \sqrt{\Delta} \partial_r + \frac{2r - r_+ - r_-}{4\pi\ell T_R \sqrt{\Delta}} \partial_\phi \right],$$

(15)

$$\bar{H}_0 = \frac{2i\nu}{\nu^2 + 3} \partial_t,$$

$$\bar{H}_1 = i e^{-\frac{\nu^2+3}{2\nu}t - 2\pi\ell T_L \phi} \left[ \frac{\nu((\nu^2 + 3)(2r - r_+ - r_-) + 8\pi\ell T_L)}{(\nu^2 + 3)^2 \sqrt{\Delta}} \partial_t \right. \\ \left. + \sqrt{\Delta} \partial_r - \frac{2}{(\nu^2 + 3) \sqrt{\Delta}} \partial_\phi \right],$$

$$\bar{H}_{-1} = i e^{\frac{\nu^2+3}{2\nu}t + 2\pi\ell T_L \phi} \left[ \frac{\nu((\nu^2 + 3)(2r - r_+ - r_-) + 8\pi\ell T_L)}{(\nu^2 + 3)^2 \sqrt{\Delta}} \partial_t \right. \\ \left. - \sqrt{\Delta} \partial_r - \frac{2}{(\nu^2 + 3) \sqrt{\Delta}} \partial_\phi \right],$$

• where

$$T_R = \frac{(\nu^2 + 3)(r_+ - r_-)}{8\pi\ell}, \quad (17)$$

$$T_L = \frac{\nu^2 + 3}{8\pi\ell} \left( r_+ + r_- - \frac{\sqrt{r_+ r_- (\nu^2 + 3)}}{\nu} \right), \quad (18)$$

$$\Delta = (r - r_+)(r - r_-). \quad (19)$$

• It is not difficult to see that the generators (15) and (16) satisfy the  $SL(2, R) \times SL(2, R)$  algebra

$$\begin{aligned} [H_n, H_m] &= i(n - m)H_{n+m} \\ [\bar{H}_n, \bar{H}_m] &= i(n - m)\bar{H}_{n+m} \\ [H_n, \bar{H}_m] &= 0 \end{aligned} \quad (n, m = -1, 0, 1) \quad (20)$$

- Hence the scalar Laplacian can be written as the  $SL(2, R)$  Casimir and the  $SL(2, R)_L \times SL(2, R)_R$  weights of the scalar field are

$$(h_L, h_R) = \left( \frac{1}{2} \sqrt{1 + \frac{4\ell^2 m^2}{\nu^2 + 3}} - \frac{1}{2}, \frac{1}{2} \sqrt{1 + \frac{4\ell^2 m^2}{\nu^2 + 3}} - \frac{1}{2} \right). \quad (21)$$

- Generators of hidden  $SL(2, R) \times SL(2, R)$  are not periodic under  $\phi \sim \phi + 2\pi$ . Periodic identification of  $\phi$  breaks the hidden symmetry to  $U(1) \times U(1)$ .
- This identification is generated by the group element

$$e^{\partial_\phi} = e^{-i2\pi\ell(T_R H_0 + T_L \bar{H}_0)} \quad (22)$$

- Interestingly, the identification (22) can be used to produce the warped black holes as a quotient of the warped  $AdS_3$  and  $T_R$  and  $T_L$  may be interpreted as the right and left temperatures of the conformal field theory dual to the warped  $AdS$  black holes.
- The Entropy calculation using the Cardy formula

$$S = \frac{\pi^2 \ell}{3} (c_L T_L + c_R T_R) \quad (23)$$

supports this hypothesis.



# Absorption cross section

- After imposing condition (13), the absorption cross section of the scalar field takes the following form

$$\begin{aligned} \sigma_{abs} \sim & \sinh \left[ 2\pi (\Omega_+ \omega + Uk) \right] \left| \Gamma \left( \frac{1}{2} + \frac{1}{2} \sqrt{1 + \frac{4m^2 \ell^2}{\nu^2 + 3}} \right. \right. \\ & \left. \left. - i(\omega(\Omega_+ + \Omega_-) + 2Uk) \right) \right|^2 \\ & \times \left| \Gamma \left( \frac{1}{2} + \frac{1}{2} \sqrt{1 + \frac{4m^2 \ell^2}{\nu^2 + 3}} - i\omega(\Omega_+ - \Omega_-) \right) \right|^2 \end{aligned} \quad (24)$$

where

$$U = \frac{2}{(r_+ - r_-)(\nu^2 + 3)} \quad (25)$$

$$\Omega_+ = \frac{2\nu r_+ - \sqrt{r_+ r_-}(\nu^2 + 3)}{(r_+ - r_-)(\nu^2 + 3)} \quad (26)$$

$$\Omega_- = \frac{2\nu r_- - \sqrt{r_+ r_-}(\nu^2 + 3)}{(r_+ - r_-)(\nu^2 + 3)} \quad (27)$$

● By defining

$$\omega_L \equiv \delta E_L = \frac{1}{2\ell} \left( \nu(r_+ + r_-) - \sqrt{r_+ r_-}(\nu^2 + 3) \right) \omega$$

$$\omega_R \equiv \delta E_R = \frac{1}{2\ell} \left( \nu(r_+ + r_-) - \sqrt{r_+ r_-}(\nu^2 + 3) \right) \omega + \frac{k}{\ell}$$

which are the left and right conjugate charges, we can rewrite (24) as

$$\sigma_{\text{abs}} \sim T_L^{2h_L-1} T_R^{2h_R-1} \sinh \left( \frac{\omega_L}{2T_L} + \frac{\omega_R}{2T_R} \right) \left| \Gamma \left( h_L + i \frac{\omega_L}{2\pi T_L} \right) \right|^2 \left| \Gamma \left( h_R + i \frac{\omega_R}{2\pi T_R} \right) \right|^2. \quad (28)$$

which is the well-known absorption cross section for a 2d CFT.